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# COMPARISON OF THE PERFORMANCES OF THE WIDE-BAND BASED WEIGHTED-SUM-OF-GRAY-GASES AND BOX MODELS IN THE CALCULATION OF NON-GRAY BOUNDARIES AT HIGH PRESSURES

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**Abstract.** *Thermal radiation is a mechanism of heat transfer that generally dominates the heat exchange in flames, furnaces and fires at high temperatures. In combustion problems, a difficulty is that typical products of the burning process (water vapor, carbon dioxide and soot) participate in the thermal exchange, absorbing, emitting or scattering radiation. In most applications, the boundaries are treated as black, which allows the decoupling of the spectral integration of the medium in relation to the surface, so that the gas models can be developed independently of the conditions at the boundary. Alternatively to the spectral integration of the radiative transfer equation (RTE) through the line-by-line (LBL) method, the spectral models reduce the computational cost, keeping the accuracy of the solution. In this paper, the wide-band based weighted-sum-of-gray-gases (WBW) model solves the radiative transfer of a participating medium with a  $H_2O-CO_2$  mixture confined by non-gray walls. In this model, the spectrum is divided into bands, and the standard weighted-sum-of-gray-gases (WSGG) model is used to solve each segment. The total pressure of the system assumes values between 1 atm and 20 atm. Results show that the WBW model has a higher accuracy than conventional box models, with maximum errors of 4%.*

**Keywords:** *thermal radiation, wide-band models, box models, non-gray surfaces, high pressures*

## 1. INTRODUCTION

Recent efforts to decrease their harmful effects on the environment have made applications for combustion at high pressure scenarios quite popular. Rockets, gas turbines, and piston engines are only a few examples of combustion devices that work at pressures higher than atmospheric (Modest and Haworth, 2016). The use of oxy-combustion for carbon capture and storage is a further use of high pressure combustion, with the primary objective of reducing atmospheric  $CO_2$  emissions (Wall *et al.*, 2011). A few of the combustion byproducts are also participating gases, such water vapor and carbon dioxide, which can emit, absorb, or scatter thermal radiation. In this framework, it is crucial to understand how to accurately calculate the radiative heat transfer in participating medium in order to enhance the combustion processes, especially the systems that operate at high pressures, which are more expensive than those at atmospheric pressure.

The total pressure of the system influences on thermal radiation, which is often the main mechanism of heat transfer in combustion processes due to the high temperatures involved. This importance becomes even greater for high pressures, since the radiation absorption increases with the total pressure, which, in turn, also affects the chemical reactions (Chen *et al.*, 2007). Soot radiation also affects the chemical reactions and has a significant dependence on the pressure, mainly through increases in the soot formation (Modest and Haworth, 2016).

The radiative transfer equation (RTE) is the mathematical expression that governs the heat transfer by thermal radiation in a participating medium. The existence of high-resolution spectral databases, such as HITEMP (high-temperature molecular spectroscopic database) (Rothman *et al.*, 2010), which offer details on the absorption coefficient of various chemical species, makes it possible to achieve the RTE solution through line-by-line (LBL) integration (Taine, 1983), despite the fact that it requires a long calculation process. Alternatively to the LBL integration, there is the weighted sum-of-gray-gases (WSGG) model (Hottel and Sarofim, 1967), which replaces the highly irregular behavior of the absorption coefficient of a gas by some equivalent gray gases with uniform absorption coefficients plus transparent windows. Each gray gas is represented by a weighting factor, which, together with the absorption coefficient, are determined by fitting total emittance data calculated through the radiative properties present in the databases. The WSGG model has generally worked effectively under atmospheric pressure circumstances (Dorigon *et al.*, 2013; Bordbar *et al.*, 2014; Selhorst *et al.*, 2022), but there are few studies using coefficients generated through emittance data at elevated pressures (Coelho and França, 2018; Bordbar *et al.*, 2021; Fonseca *et al.*, 2023b). There are some studies in which the researchers have attempted to apply these parameters to high pressure scenarios due to the dearth of WSGG coefficients at pressures above the atmospheric level, but their efforts

have often failed (Kez *et al.*, 2016; Chu *et al.*, 2017). Other spectral models, including the spectral-line based WSGG (SLW), statistical narrow-band (SNB), statistical narrow-band correlated- $k$  (SNBCK), and full-spectrum correlated- $k$  (FSCK) models, were also extended to high pressure scenarios. The results were found to be satisfactorily accurate when compared to the benchmark solution (Chu *et al.*, 2016, 2017).

In the spectral modeling field, the boundaries of the domain are typically represented as black surfaces. In order to develop spectral gas models regardless of the characteristics at the boundaries, this assumption enables the decoupling of the spectral integration of the medium in respect to the surface. The investigation of radiative exchanges between participating gases and non-gray surfaces has received minimal attention in the literature, in contrast to research in which the boundaries are black. In order to take into consideration the reflection of the radiation intensity on the walls, which is given as a function of the intensity incident on them (which, in turn, depends on the properties of the medium), the RTE must be solved iteratively at the non-gray boundaries. Due to the fact that the intensities at the boundary are known on black surfaces and the RTE may thus be solved in a single step, this feature makes the calculation computationally more expensive. Denison and Webb (1994), Solovjov *et al.* (2013), Fonseca *et al.* (2018), Liu *et al.* (2019), Fonseca *et al.* (2021) and Fonseca *et al.* (2022) are some of the authors who proposed formulations for the treatment of non-gray surfaces in the context of the spectral models.

In the present paper, the spectral integration of the RTE is solved using the wide-band based weighted-sum-of-gray-gases (WBW) model (Fonseca *et al.*, 2023a). In order to enhance the performance of the standard WSGG model, the WBW model seeks to refine the spectrum by segmenting it into a number of spectral bands and using it to solve each one of these intervals. The proposed method is used to solve the radiative heat transfer between a participating medium—composed of a  $H_2O$ - $CO_2$  mixture—and the non-gray walls that delimit the domain. The total pressure of the one-dimensional system is varied between the values of 1 atm and 20 atm. It is the first time that the effects of non-gray boundaries and elevated pressures are combined and solved through the WBW model. The accuracy of the results obtained with the proposed methodology are compared with the LBL integration and with a box model, which uses a gray approach to describe the radiative properties of the participating gases.

## 2. RADIATION MODELING

In this section, the concepts and theoretical formulation that govern the heat transfer by thermal radiation are briefly presented.

### 2.1 The line-by-line method

The radiative transfer equation (RTE) is a first-order integral-differential equation with respect to spectral intensity that governs the heat transfer by thermal radiation, and its solution depends on the spectral intensity at a specific location in the domain. Neglecting the scattering effects, the RTE for a participating medium, in a given path  $s$ , is written as (Modest, 2013; Howell *et al.*, 2016)

$$\frac{dI_\eta}{ds} = -\kappa_\eta I_\eta + \kappa_\eta I_{b\eta}. \quad (1)$$

In the above equation,  $I_\eta$  is the spectral radiation intensity,  $\kappa_\eta$  is the spectral absorption coefficient of the participating medium and  $I_{b\eta}$  is the blackbody spectral radiation intensity, which is computed from the Planck's distribution function for a given temperature  $T$  and wavenumber  $\eta$ .

Assuming a diffuse and non-gray wall at  $x = 0$ , the boundary condition associated to Eq. (1) is given by

$$I_\eta \Big|_{x=0} = I_{\eta,0} = \varepsilon_\eta I_{b\eta,0} + (1 - \alpha_\eta) \overline{I_{\eta,i}}, \quad (2)$$

where  $\varepsilon_\eta$  and  $\alpha_\eta$  are the spectral emissivity and the spectral absorptivity of the surface, respectively. The terms  $I_{b\eta,0}$  and  $\overline{I_{\eta,i}}$  are the spectral blackbody intensity evaluated at the left boundary (i.e.,  $x = 0$ ) and the average spectral intensity that arrives on the wall, respectively. The subindex  $i$  in the term  $\overline{I_{\eta,i}}$  represents the incident intensity on the wall, which is computed as

$$\overline{I_{\eta,i}} = \frac{1}{\pi} \int_0^{2\pi} (I_{\eta,i} \cos \theta_i) d\omega_i, \quad (3)$$

in which  $I_{\eta,i}$  is the spectral intensity that is incident on the wall from the direction  $d\omega_i$ . For perfectly diffuse surfaces,  $\varepsilon_\eta = \alpha_\eta$ .

The total radiative intensity and the spectral intensity at the boundary can be solved integrating these quantities over the entire spectrum, such as  $I = \int_0^\infty I_\eta d\eta$  and  $I_0 = \int_0^\infty I_{\eta,0} d\eta$ . In this way, the integrated boundary condition becomes

$$I_0 = \varepsilon I_{b,0} + (1 - \alpha) \overline{I_i}, \quad (4)$$

where the blackbody radiative intensity on the left wall and the incident intensity are calculated as  $I_{b,0} = \int_0^\infty I_{b\eta,0} d\eta$  and  $\bar{I}_i = \int_0^\infty \bar{I}_{\eta,i} d\eta$ , respectively. The total emissivity,  $\varepsilon$ , and the total absorptivity,  $\alpha$ , are computed, respectively, as

$$\varepsilon = \frac{\int_0^\infty \varepsilon_\eta I_{b\eta,0} d\eta}{I_{b,0}}, \quad (5)$$

$$\alpha = \frac{\int_0^\infty \alpha_\eta \bar{I}_{\eta,i} d\eta}{\bar{I}_i}. \quad (6)$$

The radiative heat flux,  $q_r$ , and radiative heat source,  $S_r$ , at each location in the domain can be determined by integrating the RTE and solving the spectral radiation intensities for a set of directions. These quantities can be calculated by the following equations:

$$q_r(x) = \sum_{l=1}^{n_d} \int_\eta \left\{ 2\pi\mu_l\omega_l \left[ I_{\eta,l}^+(x) - I_{\eta,l}^-(x) \right] \right\} d\eta, \quad (7)$$

$$S_r(x) = \sum_{l=1}^{n_d} \int_\eta \left\{ 2\pi\kappa_\eta\omega_l \left[ I_{\eta,l}^+(x) + I_{\eta,l}^-(x) \right] - 4\pi\kappa_\eta I_{b\eta}(x) \right\} d\eta, \quad (8)$$

in which  $n_d$  is the number of directions in which the integration of the RTE is performed. The terms  $\mu_l$  and  $\omega_l$  are the direction cosine and the weight of the Gauss-Legendre quadrature, and  $I_{\eta,l}^+$  and  $I_{\eta,l}^-$  are the forward and backward spectral radiation intensities, respectively.

## 2.2 The box model

In the present paper, the full spectrum is divided into a set of discrete bands. For comparison purposes with the proposed method, a stepwise-gray model, which is a straightforward and practical wide-band spectral model, was also implemented and tested. In this framework, a very simple box model, in which the band is approximated by a rectangular box of width  $\Delta\eta_m$  and height  $\kappa_P$  was employed, which means that the first parameter represents the size of the spectral band and the second one the absorption coefficient corresponding to that segment. So, for each spectral band, an average value of absorption coefficient was computed. Problems involving non-gray radiation have been resolved using a variety of stepwise-gray spectral models (Edwards *et al.*, 1967; Modest and Sikka, 1992; Kaminski *et al.*, 1995; Mazumder and Modest, 1999).

The Planck-mean absorption coefficient, which typically characterizes the radiative heat loss from an optically thin gas, quantifies the total emission from a medium. Furthermore, the Planck-mean absorption coefficient serves as a helpful reference in a series of approximation techniques, such as the semi-gray approximation (Modest, 2013). Once the lower and upper bounds for each spectral band have been established, the Planck-mean absorption coefficient of each segment,  $\kappa_P$ , can be computed as (Modest, 2013; Howell *et al.*, 2016)

$$\kappa_P = \frac{\int_{\Delta\eta_m} \kappa_\eta I_{b\eta} d\eta}{\int_{\Delta\eta_m} I_{b\eta} d\eta}, \quad (9)$$

where  $\Delta\eta_m$  is an arbitrary band in which the radiation spectrum was divided. In this paper, one value of  $\kappa_P$  for each one of the five spectral bands was determined. The set of these five values of  $\kappa_P$  configures the solution obtained through the box model discussed in Chapter 4.

## 2.3 The wide-band based weighted-sum-of-gray-gases model

The proposed methodology consists of applying the wide-band based WSGG (WBW) model (Fonseca *et al.*, 2023a) to solve the heat exchanges between a participating medium and the non-gray boundaries that enclose this system. The WBW model is a global spectral model in which the spectrum is divided into a set of discrete bands and the standard WSGG model is used to solve the individual contribution of each segment.

Under the assumptions of the WBW model, the RTE, for an arbitrary spectral band  $m$  and a gray gas  $j$ , is given by

$$\frac{dI_{j,m}}{dx} = -\kappa_{p,j,m} p_a I_{j,m} + \kappa_{p,j,m} p_a a_{j,m} f_m I_b, \quad (10)$$

in which  $\kappa_{p,j,m}$  is the pressure-absorption coefficient,  $p_a$  is the partial pressure of the absorbing-emitting species,  $I_{j,m}$  is the partial intensity corresponding to the  $j$ -th gray gas,  $a_{j,m}$  is the temperature coefficient,  $f_m$  is the fraction of blackbody

energy emitted from the band  $\Delta\eta_m$  and  $I_b$  is the total blackbody radiation intensity. The parameters  $\kappa_{pj,m}$  and  $a_{j,m}$  can be determined by fitting LBL emittance data through the high-resolution spectral databases, such as HITEMP 2010.

The RTE of Eq. (10) is subjected to the following boundary condition:

$$I_{j,m} \Big|_{x=0} = \varepsilon_m a_{j,m,0} f_{m,0} I_{b,0} + (1 - \alpha_m) \overline{I_{j,i,m}}, \quad (11)$$

where  $\varepsilon_m$  and  $\alpha_m$  are the partial emissivity and the partial absorptivity of the surface, respectively, of the spectral band  $m$ . The insertion of the term  $f_{m,0}$  in the above equation is due to the fact that it is necessary to account for the fraction of blackbody energy that is emitted in that spectral band. Again, the subindex  $i$  represents the incident intensity for an arbitrary gray gas  $j$ . The emissivity and absorptivity for an arbitrary band  $m$  are expressed as

$$\varepsilon_m = \frac{\int_{\Delta\eta_m} \varepsilon_{\eta,m} I_{b\eta,0} d\eta}{a_{j,0} I_{b,0}}, \quad (12)$$

$$\alpha_m = \frac{\int_{\Delta\eta_m} \alpha_{\eta,m} \overline{I_{\eta,i,m}} d\eta}{\overline{I}_i}. \quad (13)$$

Finally, after solving Eq. (10), the calculation of the radiative heat flux and radiative heat source for each spectral band follows the equations below:

$$q_r(x) = \sum_{m=1}^{n_b} \sum_{l=1}^{n_d} \sum_{j=1}^J \left\{ 2\pi\mu_l \omega_l \left[ I_{j,l,m}^+(x) - I_{j,l,m}^-(x) \right] \right\}, \quad (14)$$

$$S_r(x) = \sum_{m=1}^{n_b} \sum_{l=1}^{n_d} \sum_{j=1}^J \left\{ 2\pi\kappa_{pj,m} p a \omega_l \left[ I_{j,l,m}^+(x) + I_{j,l,m}^-(x) \right] - 4\pi\kappa_{pj,m} p a a_{j,m} f_m I_b(x) \right\}, \quad (15)$$

in which  $n_b$  is the number of bands in which the spectrum was divided, and  $I_{j,l,m}^+$  and  $I_{j,l,m}^-$  are the intensities in the forward and backward directions of the  $j$ -th gray gas.

### 3. PROBLEM DESCRIPTION AND SOLUTION METHODOLOGY

The radiative heat transfer in a one-dimensional media slab is studied in this paper. The participating medium, which is composed of a non-homogeneous and non-isothermal mixture of water vapor and carbon dioxide, is confined by two diffuse and non-gray walls that make up the 1D domain. Chapter 4 provides a description of the mole fraction and temperature profiles that are assessed. The total pressure of the system assumes the values of 1 atm, 10 atm and 20 atm; the distance between the plates is  $L = 1$  m. Figure 1 depicts the domain under study. The spectral integration of the RTE was solved using the WBW model and the directional integration was completed using the discrete ordinates method (DOM) for 8 directions, in accordance with the parameters outlined by Lathrop and Carlson (1964). 200 identically sized cells make up the spatial domain. The selection and usage of these directional and spatial discretization parameters were thoroughly detailed in Fonseca *et al.* (2023a), and it was shown to be suitable for simulations under the same physical conditions as those investigated in this paper. The details about the generation of the absorption spectra of  $H_2O$  and  $CO_2$  can be found in Dorion *et al.* (2013), and Coelho and França (2018).

A five-band stepwise variation provides the spectral emissivity distribution for the non-gray walls:  $\varepsilon_\eta = 0.5$  for  $0 \text{ cm}^{-1} < \eta < 1000 \text{ cm}^{-1}$ ;  $\varepsilon_\eta = 0.6$  for  $1000 \text{ cm}^{-1} < \eta < 2600 \text{ cm}^{-1}$ ;  $\varepsilon_\eta = 0.7$  for  $2600 \text{ cm}^{-1} < \eta < 4400 \text{ cm}^{-1}$ ;  $\varepsilon_\eta = 0.8$  for  $4400 \text{ cm}^{-1} < \eta < 6000 \text{ cm}^{-1}$ ; and  $\varepsilon_\eta = 0.9$  for  $6000 \text{ cm}^{-1} < \eta < 25\,000 \text{ cm}^{-1}$ . Although the proposed methodology is not limited to this case, both walls present the same spectral emissivity profile. The WBW method and a box model were used to solve the spectral integration of the problem, and the results were compared against LBL calculations. For  $p = 1$  atm, the set of correlations obtained with the WBW model by Fonseca *et al.* (2023a) was used, with four gray gases in each spectral band. For the other values of total pressure, the WBW coefficients generated by Fonseca *et al.* (2023b) were employed, also

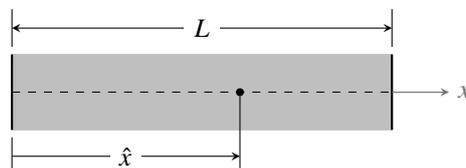


Figure 1. Representation of the one-dimensional domain.

considering that each band has four gray gases. The spectrum was segmented into the identical regions that were indicated in the description of the surfaces' emissivity profiles. The results, which were determined in terms of the radiative heat flux and radiative heat source, are presented in the following section.

#### 4. RESULTS AND DISCUSSION

The proposed methodology is tested through three different profiles of temperature and mole fraction of the participating species. The temperature profile are given by the following set of equations:

$$T(\hat{x}) = 400 + 1400 \sin^2(\pi\hat{x}), \quad (16)$$

$$T(\hat{x}) = 400 + 1400 \sin^2(2\pi\hat{x}), \quad (17)$$

$$T(\hat{x}) = \begin{cases} 880 + 920 \sin^2(2\pi\hat{x}), & \text{if } \hat{x} \leq 0.25 \\ 400 + 1400 \left\{ 1 - \sin^{3/2}[2\pi/3(\hat{x} - 0.25)] \right\}, & \text{if } \hat{x} > 0.25 \end{cases}, \quad (18)$$

where  $\hat{x} = x/L$  is a dimensionless quantity that represents the distance from the left wall. The mole fraction profiles of carbon dioxide,  $Y_c$ , follow the equations below:

$$Y_c(\hat{x}) = 0.1 \sin^2(\pi\hat{x}), \quad (19)$$

$$Y_c(\hat{x}) = 0.1 \sin^2(2\pi\hat{x}), \quad (20)$$

$$Y_c(\hat{x}) = \begin{cases} 0.1 \sin^2(2\pi\hat{x}), & \text{if } \hat{x} \leq 0.25 \\ 0.1 \left\{ 1 - \sin^{3/2}[2\pi/3(\hat{x} - 0.25)] \right\}, & \text{if } \hat{x} > 0.25 \end{cases}. \quad (21)$$

In the above equations, the maximum values of temperature and CO<sub>2</sub> mole fraction are 1800 K and 0.1, respectively; the average temperature and mole fraction are 1100 K and 0.05, respectively. Equations (16) and (19) represent profiles with simple symmetry in relation to the  $x$ -axis, with a peak located at the center of the domain, (i.e., at  $\hat{x} = 0.5$ ). Equations (17) and (20) are profiles with double symmetry in relation to the  $x$ -axis, with two peaks, at positions  $\hat{x} = 0.25$  and  $\hat{x} = 0.75$ . Equations (18) and (21) represent profiles without symmetry in relation to the  $x$ -axis, with a peak at  $\hat{x} = 0.25$ . In the three mole fraction profiles,  $Y_w = 2Y_c$ , in which  $Y_w$  is the H<sub>2</sub>O mole fraction. Figure 2 shows these profiles.

In the present paper, the type of symmetry in relation to the  $x$ -axis (or the absence of it) is combined in order to compose the test cases under study. For instance, in Case 1, the temperature and mole fraction profiles are described by the combination of Eqs. (16) and (19). Analogously, Cases 2 and 3 are described by Eqs. (17) and (20), and Eqs. (18) and (21), respectively. The cases tested here follow similar conditions to those presented in Fonseca *et al.* (2023b), but with a domain bounded by non-gray walls.

The accuracy of the test cases is evaluated through comparisons the LBL benchmark solution and the approach used to the box model. In this framework, a normalized percentage deviation was defined according to the equation below:

$$\delta\phi = \frac{|\phi_{\text{LBL}} - \phi_{\text{app}}|}{\max(|\phi_{\text{LBL}}|)}, \quad (22)$$

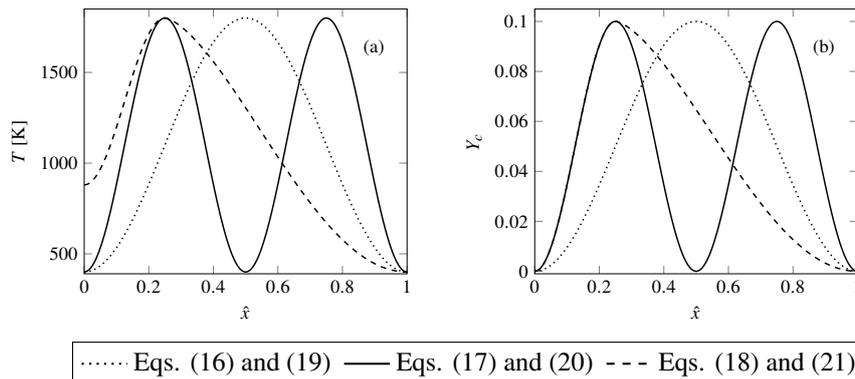


Figure 2. Profiles tested in this paper: (a) temperature profile; (b) CO<sub>2</sub> mole fraction profile.

in which  $\phi$  is either the radiative heat flux or radiative heat source, the subscripts “LBL” and “app” represent the LBL and the tested approach (the WBW or the box models in this study), and the  $\max(|\phi_{\text{LBL}}|)$  is the maximum absolute value of  $\phi$ . Later, the subscripts “max” and “avg” will be used to designate the maximum and average deviations, respectively.

Figure 3 shows the results of radiative heat flux and radiative heat source for temperature and CO<sub>2</sub> mole fraction profiles of Eqs. (16) and (19), i.e., Case 1. As it can be seen in the figure, the system is subjected to the total pressures of 1 atm, 10 atm and 20 atm, and the WBW model was compared against the LBL integration and a box model. As expected, for all the total pressure values that were tested, it was verified a better performance of the WBW model in relation to the box model. The magnitudes of these comparisons can be observed in Table 1, which shows the maximum and average percentages deviations between the tested approaches regarding the LBL benchmark solution. So, for  $p = 1$  atm, one can observe that the maximum errors in relation to the LBL solution obtained with the proposed method were 2.94 % for the radiative heat flux and 3.33 % for the radiative heat source, against 89.29 % ( $q_r$ ) and 77.3 % ( $S_r$ ) with the other approach. For  $p = 10$  atm, the maximum errors were 23.18 %, for  $S_r$  with the box model, and 2.09 %, for  $q_r$  with the WBW solution. For the pressure of 20 atm, the maximum deviations were observed for the radiative heat source for both solutions, being 17.66 % (box model) and 2.27 % (WBW model).

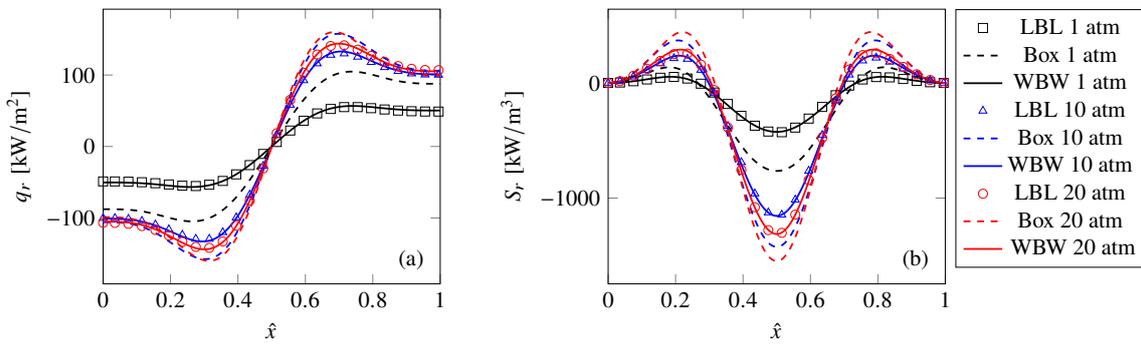


Figure 3. Results for Case 1: (a) radiative heat flux; (b) radiative heat source.

The results of  $q_r$  and  $S_r$  for Case 2 (Eqs. (17) and (20)) are depicted in Fig. 4. As in the previous case, the figure presents the curves of the results for the three values of total pressure that were investigated with the LBL, WBW and box solutions. Again comparing the two tested approaches, the WBW model was presented as the best option to solve the heat transfer, once the proposed method reached maximum deviations of 3.25 % and 3.68 % for the radiative heat flux and radiative heat source, respectively, regarding the LBL solution, for  $p = 1$  atm. The box model, in turn, led to maximum errors of 84.47 % and 73.37 % for  $q_r$  and  $S_r$ , respectively, demonstrating, again, low accuracy to describe the behavior of the solution. For  $p = 10$  atm, the maximum deviations were 22.9 % ( $q_r$ ) and 30.73 % ( $S_r$ ), with the box model, and 2.72 % ( $q_r$ ) and 2.36 % ( $S_r$ ) through the WBW model. The maximum errors obtained for the pressure of 20 atm were found for the radiative heat source with both approaches, with the magnitudes of 4.22 % and 27.81 % with WBW and box models, respectively.

Figure 5 depicts the radiative heat flux and radiative heat source for Case 3, whose temperature and CO<sub>2</sub> mole fraction profiles are given by Eqs. (16) and (19), respectively. Similarly to the previous results, in Case 3 the behavior of the solution with the WBW model presented a good match in relation to the LBL integration, with maximum deviations of 3.27 %, for  $q_r$ , and 4.35 %, for  $S_r$ , when  $p = 1$  atm. However, with the box model, the maximum errors were 84 %, for  $q_r$ , and 73.3 %, for  $S_r$ . For the other values of total pressure, the box model also presented the worst performance: for  $p = 10$  atm, the maximum errors were obtained for  $q_r$  for the box model, whose magnitude was 33.75 %, and for  $S_r$  for the WBW model, with 2.97 %; for 20 atm, these values 23 % ( $q_r$  with box model) and 2.8 % ( $S_r$  with WBW model).

For all cases evaluated in the present study, the WBW model presented a good agreement in relation to the LBL

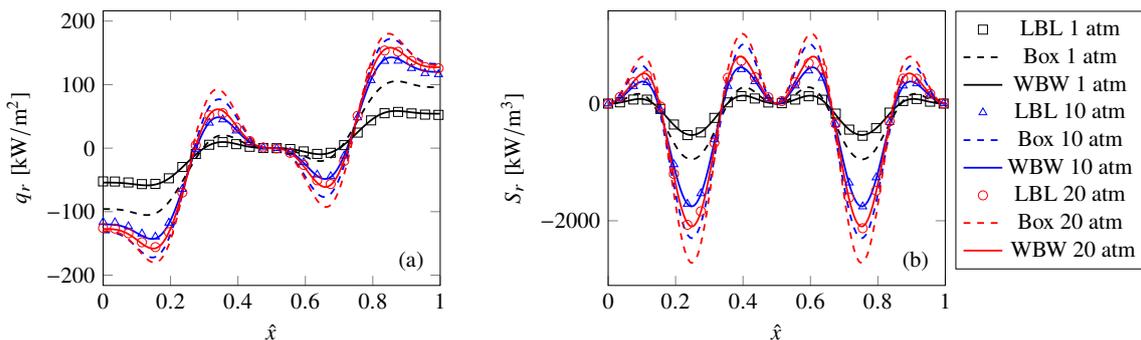


Figure 4. Results for Case 2: (a) radiative heat flux; (b) radiative heat source.

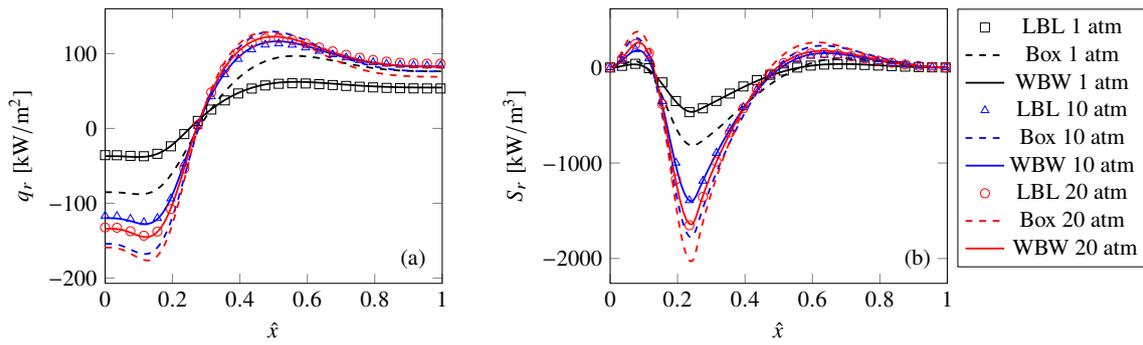


Figure 5. Results for Case 3: (a) radiative heat flux; (b) radiative heat source.

Table 1. Maximum and average normalized deviations of  $q_r$  and  $S_r$  between the LBL solution and the tested approaches.

	LBL method $\times$ Box model [%]				LBL method $\times$ WBW model [%]			
	$(\delta q_r)_{\max}$	$(\delta q_r)_{\text{avg}}$	$(\delta S_r)_{\max}$	$(\delta S_r)_{\text{avg}}$	$(\delta q_r)_{\max}$	$(\delta q_r)_{\text{avg}}$	$(\delta S_r)_{\max}$	$(\delta S_r)_{\text{avg}}$
$p = 1 \text{ atm}$								
Case 1	89.29	67.89	77.31	28.04	2.94	1.78	3.33	1.10
Case 2	84.47	41.71	73.37	27.21	3.25	1.39	3.68	1.41
Case 3	83.99	50.15	73.29	21.63	3.27	1.93	4.35	1.03
$p = 10 \text{ atm}$								
Case 1	21.32	11.88	23.18	9.43	2.09	0.97	2.03	0.81
Case 2	22.90	13.19	30.73	14.14	2.72	1.23	2.36	0.90
Case 3	33.75	12.14	27.09	6.97	2.51	1.37	2.97	0.75
$p = 20 \text{ atm}$								
Case 1	13.77	6.94	17.66	8.02	1.20	0.70	2.27	0.83
Case 2	21.97	9.77	27.81	12.43	2.55	0.88	4.22	1.32
Case 3	22.99	10.40	22.28	5.42	2.56	1.24	2.81	0.66

integration, while the box model proved inadequate to describe the radiative heat transfer. Although, with the increase in the total pressure, both models have improved their performances, the box model, in its best scenario, still presented errors of the order of 20 %, which demonstrates the effectiveness and accuracy of the proposed methodology in the sense of correctly solving the problem.

## 5. CONCLUSIONS

The WBW model was presented in this paper to describe the radiative transfer in scenarios with non-gray boundaries and pressures above the atmospheric. The WBW correlations used in the numerical simulations were generated by Fonseca *et al.* (2023a), for 1 atm, and Fonseca *et al.* (2023b), for 10 atm and 20 atm. The participating medium was composed of a non-homogeneous mixture of water vapor and carbon dioxide, confined by non-gray walls. The application of the proposed formulation led to satisfactory results regarding the LBL solution, with maximum errors of 4 %. The WBW model has also proven to be a better alternative than a gray approach typical of a box model. As potential directions for future research, an option is to evaluate more realistic emissivity profiles and to find other ways to divide the spectrum into bands without the need for them to coincide with the spectral emissivity distribution. In addition, seeking to optimize the number of gray gases in each spectral band, in order to obtain even more accurate results and with less computational effort, is also another alternative to try to improve the proposed method in future publications.

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